

Advanced Statistical Physics

Problem Set 12

due: 02.02.2011

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Problem 12.1 At what pressure the density of hydrogen will be $10\text{kg}/\text{cm}^3$? How much would this answer change if the temperature was 3000K and not 300K ? [2p]

Problem 12.2 Stellar dynamics continued. For a White Dwarf somewhat heavier than our Sun the reaction $e + p \rightarrow n + \nu$ may start. Massless and noninteracting neutrinos leave the star and the process leads effectively to the conversion of electrons and protons into neutrons. Thus a Neutron Star is born. Estimate the radius of the Neutron Star having the mass of the Sun. Use numbers from problem 11.5 [2p]

Problem 12.3 If the Neutron star is squeezed further, the neutron gas becomes relativistic. Investigate the fate of an ultrarelativistic star, where the energy of neutrons is given by

$$\varepsilon = \sqrt{m^2c^4 + p^2c^2} \approx pc. \quad (1)$$

How heavy should the initial star be to be able to give birth to the Black Hole? [2p]

Problem 12.4 Find the average squared displacement of the atom in a solid $\langle \vec{u}^2 \rangle$ in the Debye model. If you can not calculate the final integral, analyze the high/low temperature limits. The mass of the atom is m , sound velocity v , the crystal has N atoms in the volume V . [2p]

At what temperature the amplitude of the lattice vibrations will be comparable to the interatomic distance? [0p]

You start with writing the displacement of the atom i as a sum over displacements corresponding to individual phonon modes (wave vector \vec{k})

$$\vec{u}_i = \sum_k \vec{u}_i^k. \quad (2)$$

Then you notice that due to the translational symmetry the squared displacement in each mode $|\vec{u}_i^k|^2$ does not depend on the position of the atom i . Finally you know from Quantum

mechanics that the energy of an oscillator is equally split between the (average) kinetic and potential energy. Thus the thermal average $\langle \dots \rangle$ of $|\vec{u}_i^k|^2$ is related to the thermal average of the single phonon mode oscillator energy ε_k

$$\frac{Nm\omega_k^2 \langle |\vec{u}_i^k|^2 \rangle}{2} = \frac{1}{2} \langle \varepsilon_k \rangle. \quad (3)$$

Problem 12.5 Consider the (3-dimensional) Bose gas in the potential $U = m\omega^2 r^2/2$. Find the temperature of the Bose-Einstein condensation and the discontinuity of the heat capacity at this temperature. Number of particles is N . You may need functions f_m^+ and ξ_m defined in Kardar's book. [2p]